

The **SSHADÉ** project: a European Database Infrastructure in Solid Spectroscopy

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OSUG



Motivations !

- Spectroscopy and spectro-imagery are increasingly used in **space missions**, *in orbit or in situ*, to study small objects of the solar system
 - various spectroscopies used to study meteorites and cometary dusts *in the laboratory*
 - large number of laboratories in Europe have developed **experiments to study** the spectroscopic properties of a variety of solid **variety of materials (ices, minerals, organics, ...)** of astrophysical interest
 - The amount of spectral data collected is huge
 - these laboratories boast **leading-edge expertise** in some solid spectroscopy fields.
 - However **most of these data**, although published, **are very difficult to access** in a usable form (i.e. electronic) to compare with observation or to use in numerical codes.
- ➔ We decided to **extend our datamodel (SSDM) and our database (GhoSST)** in order to **build a database infrastructure able to distribute the spectroscopic data** of most of the European laboratories

What is *SSHADE*?

“Solid Spectroscopy Hosting Architecture of Databases and Expertise”

It is an **European project** of a **set of databases (~20)** to provide the community with spectra of solids (*ices, minerals, organics, cosmo-materials, ...*) of astrophysical and terrestrial interests in the **X-ray to sub-mm range**.

Consortium: SSHADE Partners

- 20 groups in 18 laboratories
- 8 countries: F, GB, E, I, CH, D, PL, HU
- Expertises in:
 - Solid materials:
 - ices, minerals
 - organics, inorganics
 - meteorites, IDPs, ...
 - optical materials, ...
 - Various spectroscopies:
 - transmission, reflection,
 - photometry, Raman, XANES ...
 - Various spectral ranges
 - X, UV, Vis, NIR-FIR, sub-mm



The SSHADE Partners: 20 (8 countries)

- **IPAG** / Planéto, Grenoble - F (Bernard Schmitt, Lydie Bonal, Damien Albert) 
- **Open University**, Milton Keynes – UK (Nigel Mason) 
- **IAS**, Univ. Paris-Sud - F (Emmanuel Dartois, Donia Baklouti) 
- **IRAP / GPPS**, Toulouse - F (Patrick Pinet, Yves Daydou) 
- **IRAP / MICMAC**, Toulouse - F (Karine Demyk, Yves Daydou) 
- **LPG**, Univ. Nantes - F (Yann Morizet, Manuel Giraud) 
- **Space and Planetary Science Division**, Univ. of Bern - CH (Antoine Pommerol) 
- **PIIM**, Univ. Aix-Marseille - F (Patrice Theulé) 
- **IAPS**, INAF, Roma - I (Alessandra Rotundi, Vincenzo della Corte) 
- **IAPS**, INAF, Roma - I (Fabrizio Capaccioni, Cristian Carli) 
- **LISA** / Univ. Paris-Est - F (Nicolas Fray) 
- **AIU Observatory**, Jena - D (Harald Mutschke, Jürgen Weiprecht) [DOCCD 'database'] 
- **Centro de Astrobiología**, INTA-CSIC – E (Guillermo Muñoz Caro) 
- **Instituto de Estructura de la Materia**, Madrid – E (Vicente Timón, M. Angel Moreno) 
- **LATMOS / IMPEC**, Institut Pierre Simon Laplace - F (Nathalie Carrasco) 
- **LGL-TPE / ENS-Lyon** - F (Bruno Reynard, Gilles Montagnac (exp.), Razvan Caracas (th.)) 
- **Konkoly Astronomical Institute** – HU (Akos Kereszturi) 
- **Planetary Geology Laboratory**, PAS – PL (Joanna Gurgurewicz, Luigi Castaldo) 
- **Clay Minerals Laboratory**, PAS– PL (Arkadiusz Derkowski, Luigi Castaldo) 
- **ESRF / FAME line**, Grenoble – EU / F (Denis Testemale, Isabelle Kieffer)  

Data of SSHADE

- **Spectral ranges:**
 - from X to sub-mm
- **Solids (+ liquids):**
 - Ices (low to high pressure, low to room temperature, mixtures, matrix isolated, ...)
 - clathrates hydrates, hydrates
 - minerals (naturals and synthesized), rocks
 - organic matter (natural and synthesized), polymers, VUV
 - Extraterrestrial matter: meteorites, IDPs, ...
 - also liquids (organics, ...)
- **Data types:**
 - **Spectra**
 - infrared transmission spectra, absorption coefficients, optical constants
 - Raman spectra et micro-spectroscopy
 - reflectance spectra of surfaces, spectro-photometric functions
 - multispectrale and multi-angular Imagery of surfaces
 - **Bandlist**
 - position, width, intensity, attribution ... for molecular solids

SSHAD E Solid Spectroscopy Data (1)

- **IPAG** (*GhoSST*)
 - Vis-FIR spectra, optical constants and bandlists of ices, minerals/rocks, organic molecules and materials, optical materials.
 - Vis-NIR bidirectionnal reflectance spectra + BRDF of surfaces: snows, ices, minerals/rocks, organic materials, salts, sulfur, ...
 - UV-Vis Raman + Fluorescence spectra + Band parameters of organic molecules, natural et synthetic carbonaceous materials, meteorites, IDPs, Stardust grains.
 - NIR+MIR microscopic spectro-images (spectral and composition maps) of natural et synthetic carbonaceous materials, minerals/rocks.
- **IAS**
 - MIR spectra of ices and clathrates hydrates
 - MIR spectra organic molecules and materials synthesized by VUV irradiation
 - MIR+FIR spectra of synthetic carbonaceous materials.
 - MIR+FIR spectra and Raman micro-spectroscopy of meteorites, IDPs, ...
- **LPGNantes**
 - Raman spectra of ices and clathrates hydrates at high pressures
 - NIR reflectance spectra of ices, clathrates hydrates and minerals
- **IRAP**
 - multispectral and multi-angular imagery of mineral surfaces
 - multispectral photometric fonctions
 - FIR+sub-mm spectra and absorption coefficients of synthetic silicates
- **LISA**
 - MIR spectra of ices and organic polymers
- **LATMOS**
 - MIR spectra of synthetic organics (tholins, ...)

SSHADe Solid Spectroscopy Data (2)

- **PIIM**
 - MIR spectra of ices and organic molecules synthesized by VUV irradiation and chemical reactions
 - MIR spectra de molecules isolated in matrices.
- **LGL-TPE, ENS-Lyon**
 - Experimental and theoretical Raman spectra of minerals and meteorites
- **Open University**
 - VUV spectra of ices
- **IAPS**
 - Spectra of meteorites, minerals and rocks
- **Univ. of Bern**
 - multi-bands bidirectional photometry of various materials: ices, organics, mixtures, ...
- **AIU Observatory** [DOCCD 'database']
 - optical constants of minerals
- **Centro de Astrobiología**
 - NIR-FIR spectra of ices ...
- **Instituto de Estructura de la Materia**
 - NIR-FIR spectra + optical constants of ices, organic molecules, phyllosilicates
- **Konkoly Astronomical Institute**
 - MIR spectra of meteorites and clay minerals
- **Planetary Geology Laboratory**
 - NIR spectra of basalts and organics
- **Clay Minerals Laboratory**
 - NIR+MIR spectra of well characterized clay minerals

What is *SSHADE*?

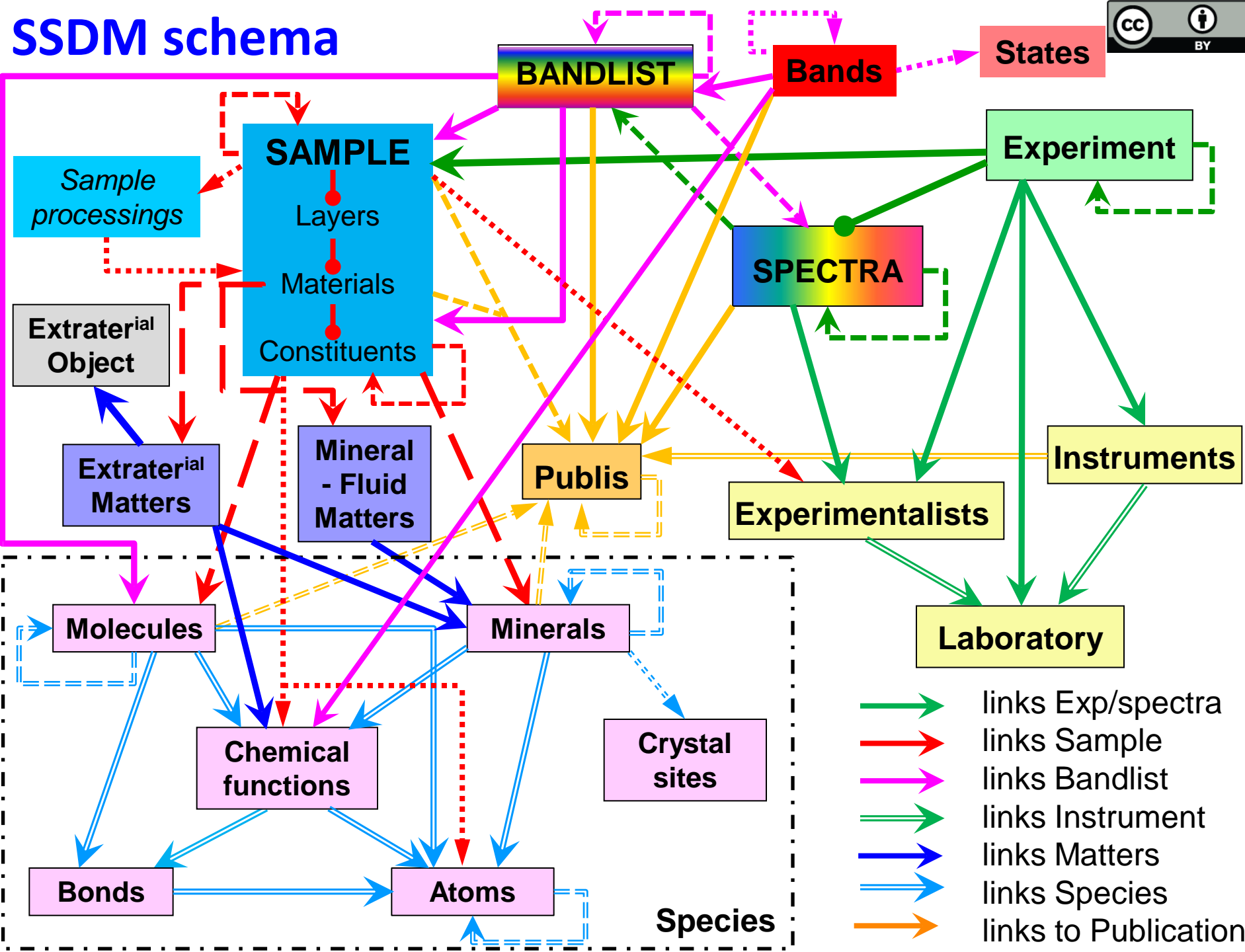
“Solid Spectroscopy Hosting Architecture of Databases and Expertise”

It is a set of databases (~20) based on
the **Solid Spectroscopy Data Model (SSDM)**
and building on the **GhoSST database** interface

Gho  DM
T
HADE


GhoSST
SSHADE

SSDM schema



Search(s) Spectra + Bandlists



View results | 6 spectra found

Material

Species type **Molecules**
Molecule
Molecular ion

Species formula Species symbol, stoichiometric formula or structural formula

Species name and code

Species relevance **Molecules**
Actual
Precursor

Matter origin Terrestrial
Extraterrestrial
Synthetic

Material name

Matter family Molecular
Mineral
Rock

Temperature min. (K)

Temperature max. (K)

Layers number

Layer texture Loose granular
Cemented granular
Sintered granular

Constituents mixing Single phase
Multi-phases
Coated grain

Constituent family Molecular solid
Molecular liquid
Molecular gas

Phase type Crystalline
Amorphous
Glassy

Species compound Pure
Mixed
Binary

Ad/absorption Adsorption
Insertion
No

Instrument

Type FTIR spectrometer
Grating spectrometer
AOTF spectrometer

Technique Transmission
Reflection
Biconical reflection

Spectral range

Unit

Type FIR
MIR
NIR

Custom min.

Custom max.

Spectrum

Type Bidirectional reflectance

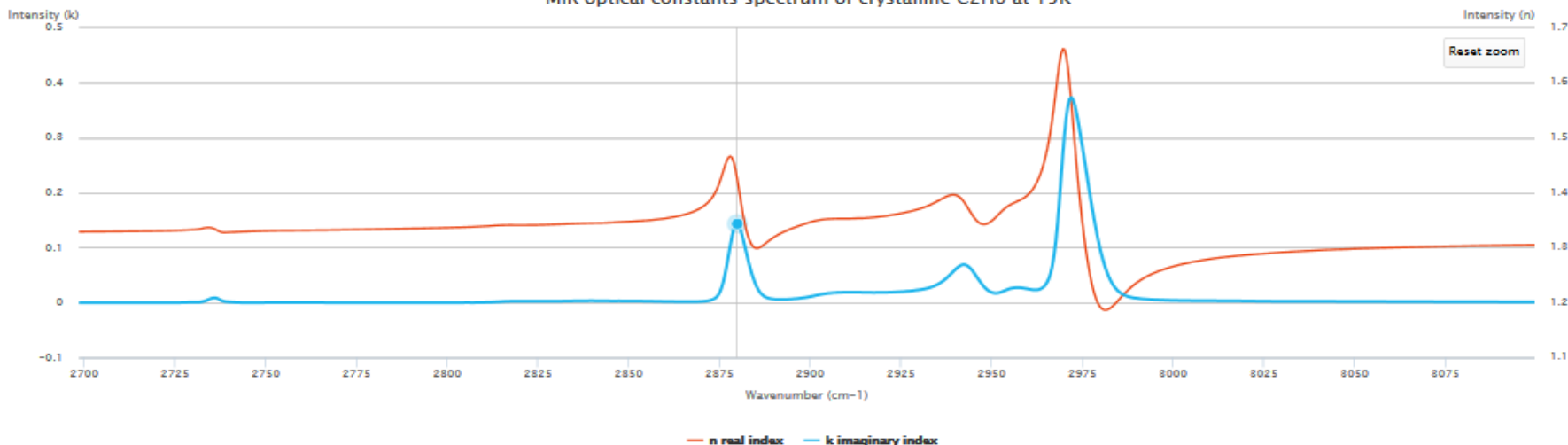
View Spectra + Metadata

Graph

Export this spectrum

Compare this spectrum

MIR optical constants spectrum of crystalline C2H6 at 15K



Point details

Wavenumber (cm-1) 2879.886
k imaginary index 0.1436379

Settings

Display Errors: Infos:
Unit cm-1
Scale X: Linear Y left: Linear Y right: Linear
Colors K: 1CB5ED N: EB4513
Zoom Full range Valid range

Spectrum

ID 474
UID SPECTRUM_BS_20130205_003
Spectrum type optical constants
Title MIR optical constants spectrum of crystalline C2H6 at 15K
Valid range 670.00-3660.00 cm^{-1}
Created 1996-01-01
Last updated 2014-08-15
Analysis Iterative inversion of n and k with full optical model of the film+substrate + Kramers-Kronig analysis over 670-3660 cm^{-1} range and $n(15800cm^{-1}, CH_4) = 1.36$ (F. Trotta thesis 1996) - k is set to 0 and n is smoothed outside C2H6 bands to reduce noise and the CO2 impurity band (2357.5 cm^{-1}) is removed by interpolation
Quality flag 4

[More details...](#)

Sample

ID 305
UID SAMPLE_BS_20130205_003
Name C2H6 crystalline - dep 15K
Number of layers 1





Spectra

Actions	ID	Type	File title	Spectral range min	Spectral range max	Sample temperature	Species	Date
View <input type="checkbox"/>	10	transmission	N87_S09 CH4 ICE 30 K	1850	10500	-	-	2011-08-04
View <input checked="" type="checkbox"/>	94	bidirectional reflectance	NIR-MIR bidirectional reflection spectrum ($i=0^\circ$, $e=30^\circ$) of Smectite SWy-2 with adsorbed H2O at -30°C an P(H2O)= 0.25 mbar	1.2	2	-	-	2011-10-01

Wavenumber/Wavelength/Frequency

Unit:

Format:

Number of significant digits:

Number of decimals:

Value/Intensity

Format:

Number of significant digits:

Number of decimals:

Spectra data file

Type:

Format:

Export archive file

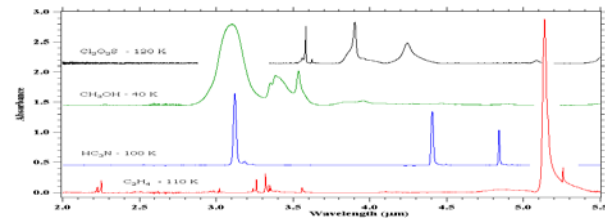
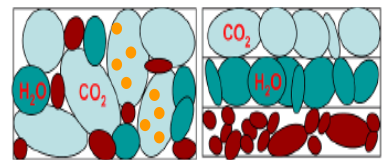
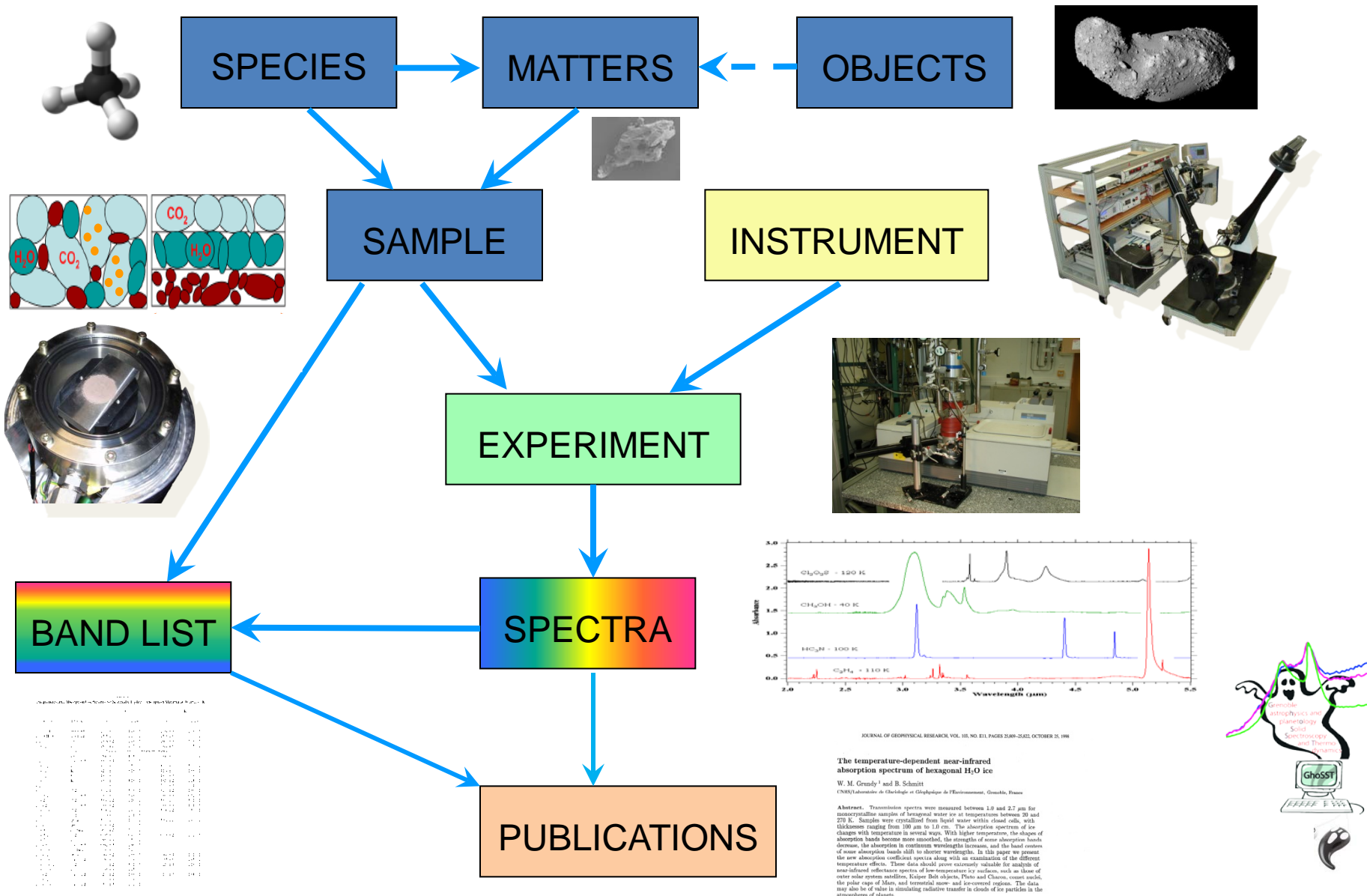
Filename:

Compression type:

Hosting Solid Spectroscopy data of European Data Providers: *SSHADÉ*

- Based on the *GhoSST* database developments (Europlanet + VAMDC)
=> <http://ghosst.osug.fr>
- Made of:
 - ✓ A 'solid spectroscopy' interface
 - ✓ A Search/Visualization/Export engine
 - ✓ A set of databases: one per group (GhoSST is one of them)
 - ✓ A common fundamental database
 - ✓ A common bandlist database
- Also direct link from each individual DB web page (powered by SSHADÉ)
- All databases/engines hosted at OSUG data center (OSU Grenoble - UJF)
- SSHADÉ will be a service of others VO (VESPA (*Europlanet-VO*), VAMDC, ...)

GhoSST: SSDM General Structure



Wavelength (µm)	Wavenumber (cm⁻¹)	Band Name
2.13	4695	ν₁
2.15	4651	ν₁
2.17	4607	ν₁
2.19	4563	ν₁
2.21	4519	ν₁
2.23	4475	ν₁
2.25	4431	ν₁
2.27	4387	ν₁
2.29	4343	ν₁
2.31	4299	ν₁
2.33	4255	ν₁
2.35	4211	ν₁
2.37	4167	ν₁
2.39	4123	ν₁
2.41	4079	ν₁
2.43	4035	ν₁
2.45	3991	ν₁
2.47	3947	ν₁
2.49	3903	ν₁
2.51	3859	ν₁
2.53	3815	ν₁
2.55	3771	ν₁
2.57	3727	ν₁
2.59	3683	ν₁
2.61	3639	ν₁
2.63	3595	ν₁
2.65	3551	ν₁
2.67	3507	ν₁
2.69	3463	ν₁
2.71	3419	ν₁
2.73	3375	ν₁
2.75	3331	ν₁
2.77	3287	ν₁
2.79	3243	ν₁
2.81	3199	ν₁
2.83	3155	ν₁
2.85	3111	ν₁
2.87	3067	ν₁
2.89	3023	ν₁
2.91	2979	ν₁
2.93	2935	ν₁
2.95	2891	ν₁
2.97	2847	ν₁
2.99	2803	ν₁
3.01	2759	ν₁
3.03	2715	ν₁
3.05	2671	ν₁
3.07	2627	ν₁
3.09	2583	ν₁
3.11	2539	ν₁
3.13	2495	ν₁
3.15	2451	ν₁
3.17	2407	ν₁
3.19	2363	ν₁
3.21	2319	ν₁
3.23	2275	ν₁
3.25	2231	ν₁
3.27	2187	ν₁
3.29	2143	ν₁
3.31	2099	ν₁
3.33	2055	ν₁
3.35	2011	ν₁
3.37	1967	ν₁
3.39	1923	ν₁
3.41	1879	ν₁
3.43	1835	ν₁
3.45	1791	ν₁
3.47	1747	ν₁
3.49	1703	ν₁
3.51	1659	ν₁
3.53	1615	ν₁
3.55	1571	ν₁
3.57	1527	ν₁
3.59	1483	ν₁
3.61	1439	ν₁
3.63	1395	ν₁
3.65	1351	ν₁
3.67	1307	ν₁
3.69	1263	ν₁
3.71	1219	ν₁
3.73	1175	ν₁
3.75	1131	ν₁
3.77	1087	ν₁
3.79	1043	ν₁
3.81	999	ν₁
3.83	955	ν₁
3.85	911	ν₁
3.87	867	ν₁
3.89	823	ν₁
3.91	779	ν₁
3.93	735	ν₁
3.95	691	ν₁
3.97	647	ν₁
3.99	603	ν₁
4.01	559	ν₁
4.03	515	ν₁
4.05	471	ν₁
4.07	427	ν₁
4.09	383	ν₁
4.11	339	ν₁
4.13	295	ν₁
4.15	251	ν₁
4.17	207	ν₁
4.19	163	ν₁
4.21	119	ν₁
4.23	75	ν₁
4.25	31	ν₁

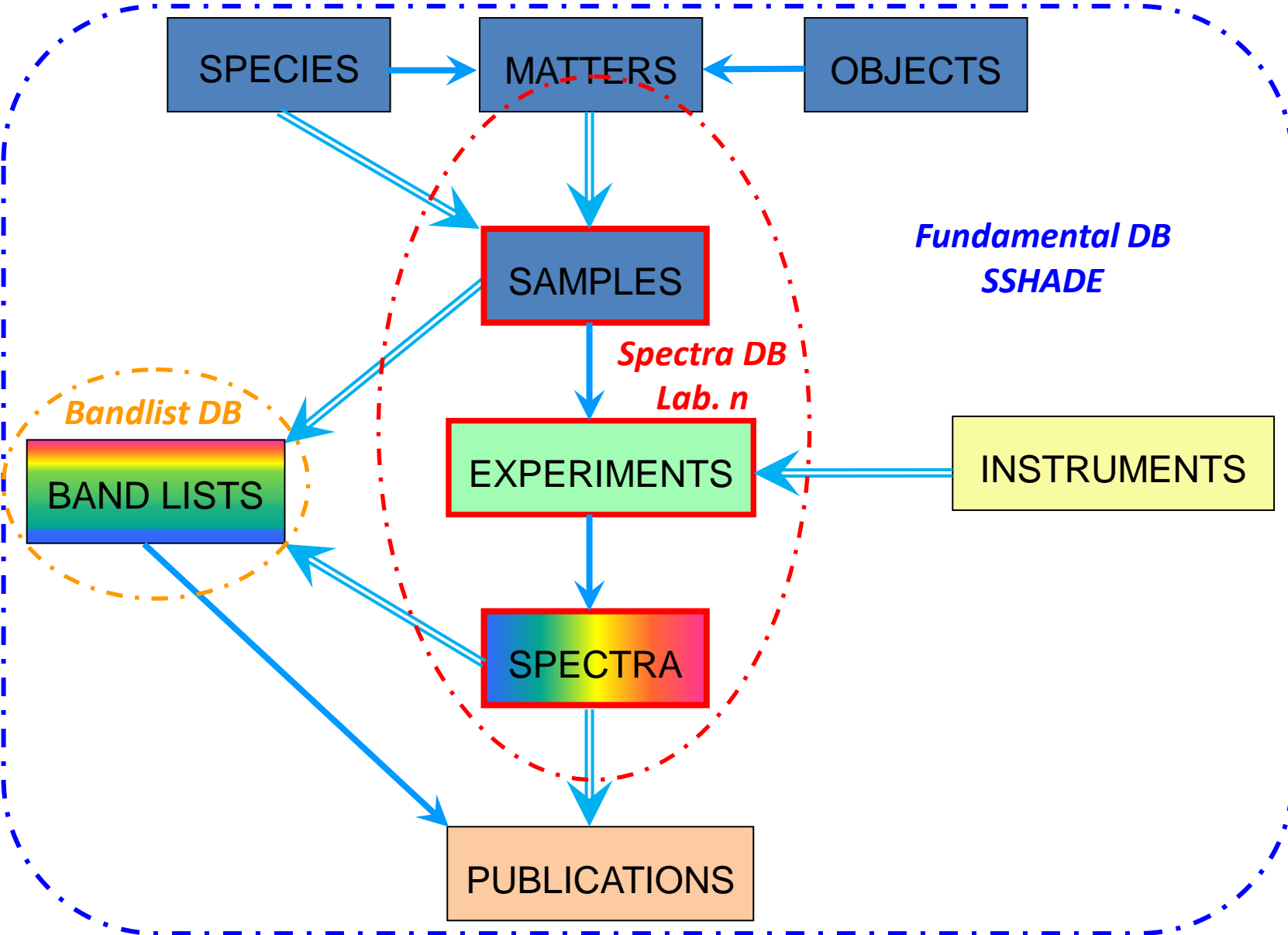
JOURNAL OF GEOPHYSICAL RESEARCH, VOL. 103, NO. E11, PAGES 2489-2502, OCTOBER 25, 1998

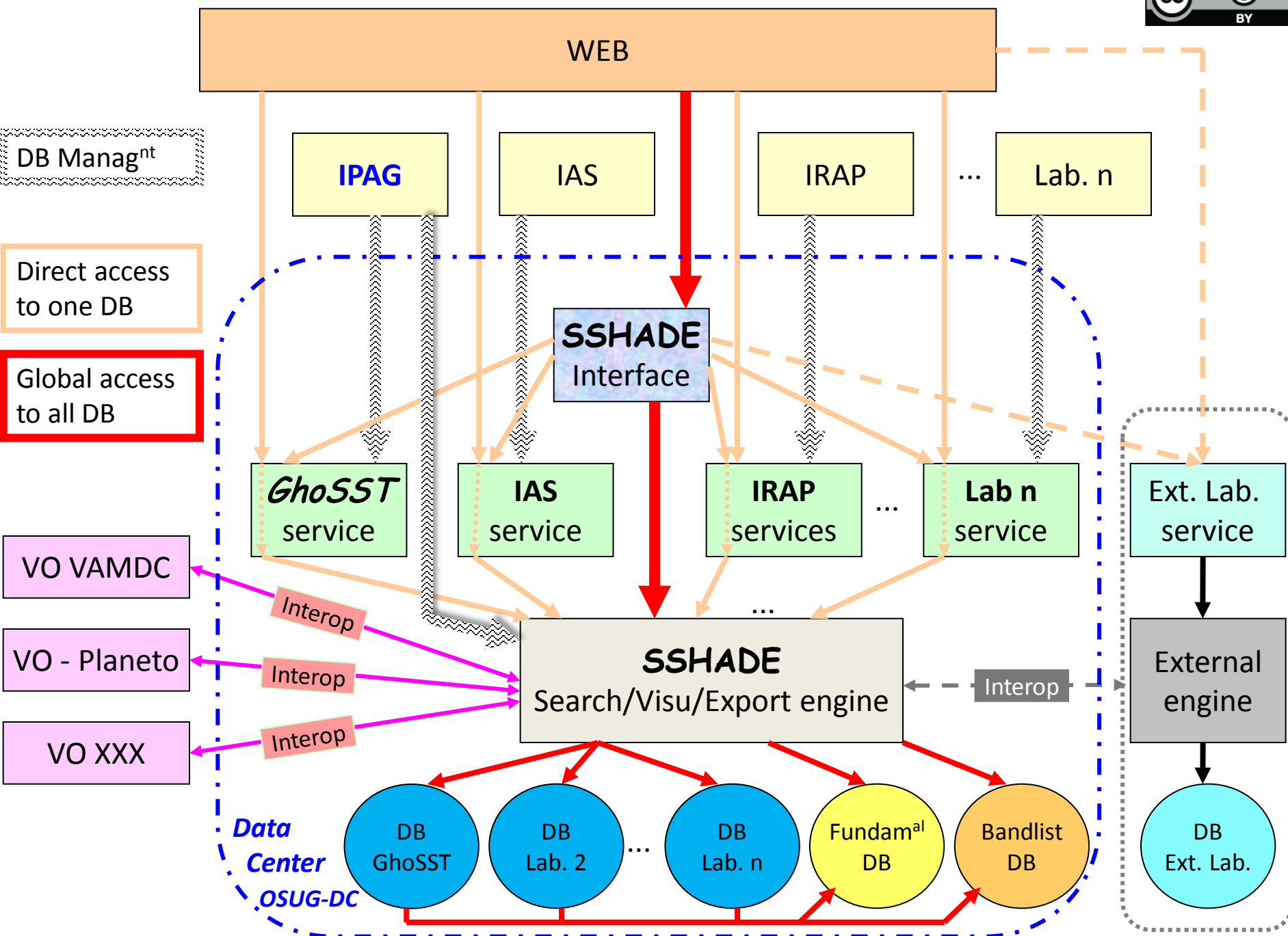
The temperature-dependent near-infrared absorption spectrum of hexagonal H₂O ice
 W. M. Grundy¹ and B. Schmitt
 CNRS/Laboratoire de Chimie et de Physique de l'Environnement, Grenoble, France

Abstract. Transmission spectra were measured between 2.0 and 2.7 μm for microcrystalline samples of hexagonal water ice at temperatures between 20 and 270 K. Samples were crystallized from liquid water within closed cells, with thicknesses ranging from 100 μm to 1.0 cm. The absorption spectrum of ice changes with temperature in several ways. With higher temperature, the shapes of absorption bands become more smoothed, the strengths of some absorption bands decrease, the absorption in continuum wavelengths increases, and the band centers of some absorption bands shift to shorter wavelengths. In this paper we present the near-infrared reflectance spectra of low-temperature ice samples, only on those of outer solar system satellites, Kuiper Belt objects, Pluto and Charon, comets nuclei, the polar caps of Mars, and interplanetary dust and cometary regions. The data may also be of value in simulating radiative transfer in clouds of ice particles in the atmosphere of planets.



SSHADE: new SSDM Structure





SSHADÉ Project

- **Done:**

- Feb. 2014 : **first SSHADÉ-Europe meeting**
- mid-2014: start 2-3 new DataBases (in pre-SSHADÉ structure)
- sept. 2014 Europlanet-RI proposal (*Horizon 2020*): VESPA WPs

- **Project:**

- mid-2015/17: **develop SSHADÉ infrastructure**
(if Horizon 2020 funding)
- 2015/19 prepare data for ingestion (in pre-SSHADÉ structure)
- 2017: ingest all data in SSHADÉ database
- 2017: **open first SSHADÉ DataBases (~6-8) to ‘users’**
- 2019: **≥ 18 databases open to ‘users’**

to follow SSHADÉ:

<http://blog.sshade.eu/>